Genealogy and spatial distribution of the N-particle branching random walk with polynomial tails

> Sarah Penington University of Bath

Joint work with Matt Roberts and Zsófia Talyigás

Branching-selection systems

- Particle systems: particles branch (produce offspring) and move in space killing rule keeps total number of particles constant.
- · Toy models for a population under selection.
 - Location of a particle (= individual) represents its evolutionary fitness.
- Introduced by Brunet and Derrida in 1990s.
 Recent results and open conjectures about long-term behaviour.
 Genealogy:



Let X be a real-valued random variable (jump distribution).

At each time nello, each particle has two offspring.

Each of the 2N offspring particles makes an independent jump from its parent's location, with the same law as X.

The N rightmost particles (of the 2N offspring particles) form the population at time n+1.



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Light-tailed jump distribution Asymptotic speed If $\mathbb{E}[X] < \infty$ then $\exists v_N \in (0, \infty)$ s.t. $\lim_{N \to \infty} \frac{\chi_N^{(N)}(n)}{n} = v_N = \lim_{N \to \infty} \frac{\chi_{1}^{(N)}(n)}{n}$ a.s. and $n \to \infty$ $\frac{1}{n} = v_N = \lim_{N \to \infty} \frac{\chi_{1}^{(N)}(n)}{n}$ a.s. and $n \to \infty$ $\frac{1}{n} = v_N = \lim_{N \to \infty} \frac{\chi_{1}^{(N)}(n)}{n}$ a.s. and $\frac{1}{n} = \frac{1}{n}$ <u>Theorem</u> (Bérard and Gouéré 2010) If $\mathbb{E}[e^{XX}] < \infty$ for some $\lambda > 0$ (+ technical assumptions) then $\lim_{N \to \infty} v_N = v_\infty$ exists and $v_\infty - v_N \sim c (\log N)^{-2}$ as $N \to \infty$. Conjectured by Branet + Derrida 1997. Related result for Fisher - KPP equation with noise

(Mueller, Mytnik, Quastel 2009)

Light-tailed jump distribution Asymptotic speed $\lim_{n \to \infty} \frac{X_{N}^{(N)}(n)}{n} = V_{N} = \lim_{n \to \infty} \frac{X_{1}^{(N)}(n)}{n} \quad \text{a.s. and} \quad \lim_{n \to \infty} L^{1}.$ If $\mathbb{E}[X] < \infty$ then $\exists v_N \in (0, \infty)$ s.t. Theorem (Bérard and Gouéré 2010) If $\mathbb{E}[e^{\lambda X}] < \infty$ for some $\lambda > 0$ (+technical assumptions) $\lim_{N \to \infty} V_N = V_{\infty} \text{ exists and } V_{\infty} - V_N \sim C (\log N)^{-2} \text{ as } N \to \infty.$ then Conjectured by Brunet + Derrida 1997. Related result for Fisher-KPP equation with noise (Mueller, Mytnik, Quastel 2009) Genealogy Sample k particles from the N particles and trace their ancestry backwards in time \rightarrow coalescent process. Conjecture (Brunet, Derrida, Mueller, Munier) If X is light-tailed then the genealogy of a sample on a $(\log N)^3$ timescale converges to a Bolthausen-Sznitman coalescent as $N \rightarrow \infty$.

Coalescent processes

Kingman's coalescent Neutral population: choose particles to kill uniformly at random in each generation.



Bolthausen-Sznitman coalescent

Population under selection.



Thanks to Götz Kersting and Anton Wakolbinger

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N-BRW with heavy-tailed jump distribution Suppose $\mathbb{P}(X > \infty) \sim \infty^{-\alpha}$ as $\infty \to \infty$, for some $\alpha > 0$. Asymptotic speed Theorem (Bérard and Maillard 2014) If $\mathbb{E}[X] < \infty$, $\lim_{n \to \infty} \frac{X_{N}^{(N)}(n)}{n} = V_{N}$ where $V_{N} \sim C_{\alpha} N^{\prime \prime \alpha} (\log N)^{\prime \prime \alpha - 1}$ as $N \rightarrow \infty$. If $\mathbb{E}[X] = \infty$, cloud of particles accelerates. Genealogy Conjecture (Bérard and Maillard) The genealogy on a log N timescale is approximately given by a Star-shaped coalescent when N is large.

Time and space scales

Let
$$P(X > \infty) = \frac{1}{h(\infty)}$$
 for $\infty \ge 0$.

Assume h is regularly varying with index 2>0

i.e. for any
$$y>0$$
, $\frac{h(xy)}{h(x)} \longrightarrow y^{\alpha}$ as $x \to \infty$.

and $\mathbb{P}(X \ge 0) = 1$ (no negative jumps).

Let $l_N = \lceil \log_2 N^7 \rceil$ time scale

Let $a_N = h^{-1}(2N\ell_N)$, where $h^{-1}(\infty) := \inf \{y, z, 0\} : h(y) > \infty \}$. space scale $E[\# jumps of size > c_1 a_N in a time interval of length c_2 l_N]$ $= 2N \cdot c_2 \ell_N \mathbb{P}(X > c_1 a_N)$ $= \frac{2Nc_2\ell_N}{h(c_1a_N)} \sim \frac{2Nc_2\ell_N}{c_1^{\alpha}2N\ell_N} = \frac{c_2}{c_1^{\alpha}}$

as $N \rightarrow \infty$

Main result

 $\omega.h.p. = \omega ith probability \rightarrow 1$ as $N \rightarrow \infty$. Theorem (P., Roberts, Talyigás 2021) For $\eta > 0$, kell and $t > 4\ell_N$, the following occurs $\omega.h.p.:$ • Spatial distribution: At time t, there are N - o(N) particles in $[X_{1}^{(N)}(t), X_{1}^{(N)}(t) + \eta a_{N}].$ • Genealogy: The genealogy on an l_N -timescale is asymptotically given by a star-shaped coalescent. i.e. $\exists T \in [t - 2l_N, t - l_N]$ s.t. ω .h.p., for a uniform sample of k particles at time t, every particle is descended from the rightmost particle at time T and no pair of particles in the sample has a common ancestor after time $T + \Sigma_N \ell_N$, for any $(\Sigma_N)_N$ with $\Sigma_N \rightarrow 0$ and $\Sigma_N \ell_N \rightarrow \infty$ as $N \rightarrow \infty$.



 $\exists T \in [t-2l_N, t-l_N] \text{ s.t. } \omega.h.p., \text{ for a uniform sample of } k \text{ particles} \\ \text{at time } t, \text{ every particle is descended from the rightmost particle at time } T \\ \text{and no pair of particles in the sample has a common ancestor after time} \\ T + E_N l_N, \text{ for any } (E_N)_N \text{ with } E_N \rightarrow O \text{ and } E_N l_N \rightarrow \infty \text{ as } N \rightarrow \infty. \end{cases}$

Spatial distribution

At time t, there are N - o(N) particles in $[X_1^{(N)}(t), X_1^{(N)}(t) + \eta \alpha_N]$ w.h.p. <u>Proposition</u> (PRT 2021) There exist $0 < p_r \leq q_r < 1$ s.t. $q_r \rightarrow 0$ as $r \rightarrow \infty$ and $p_r \rightarrow 1$ as $r \rightarrow 0$ s.t. for r > 0, for N sufficiently large and $t > 3l_N$, $\mathbb{P}\left(X_{N}^{(N)}(t)-X_{I}^{(N)}(t) \ge ra_{N}\right) \in [p_{r}, q_{r}].$ Genealogy ω .h.p. $\exists T \in [t - 2l_N, t - l_N]$ s.t. ω .h.p., for a uniform sample of k particles at time t, every particle is descended from the rightmost particle at time T and no pair of particles has a common ancestor after time $T + o(\ell_N)$. <u>Proposition</u> (PRT 2021) For $0 \le s_1 \le s_2 \le 1$, $\exists p > 0$ s.t. for N sufficiently large and t>4lN, $\mathbb{P}(\mathsf{T} \in [t-2\ell_N + s_1\ell_N, t-2\ell_N + s_2\ell_N]) > p.$

Light-tailed $\mathbb{P}(X > \infty) \leq e^{-cx}$, c>0

Time to coalesce Coalescent (log N)³ Bolthausen-Sznitman

Heavy-tailed $P(X > x) \sim x^{-\alpha}, \alpha > 0$

log N Star-shaped

Time to coalesceCoalescentLight-tailed
$$P(X > \infty) \le e^{-cx}$$
, c>0 $(\log N)^3$ Bolthausen-SznitmanStretched exponential $P(X > \infty) \sim e^{-\infty \beta}$, $\beta \in (0, 1)$????tail $P(X > \infty) \sim e^{-\alpha}$, $\alpha > 0$ $\log N$ Star-shaped

Work in progess with Z. Talyigás.

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Work in progess with Z. Talyigás. Simulation by Z. Talyigás.

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- Each particle, independently, branches
 into two particles after an Exp(1) time.
- Each time a particle branches, the leftmost particle in the system is killed.





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N particles in the system at all times. Introduced by Maillard (2012). Conjecture (Brunet/Derrida, Maillard): Genealogy of a sample on a $(\log N)^3$ timescale converges to a Bolthausen-Sznitman coalescent as $N \rightarrow \infty$. One tool: over a fixed timescale, as $N \rightarrow \infty$, density converges to solution of

One tool: over a fixed timescale, as $N \rightarrow \infty$, density converges to solution of a free boundary problem. (Hydrodynamic limit: De Masi/Ferrari/Presutti/Soprano-Loto '17)

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Warm up lemma Recall $l_N = \prod_{0 \ge 2} N^7$. Recall $X_1^{(N)}(t) \le X_2^{(N)}(t) \le \dots \le X_N^{(N)}(t)$ ordered particle positions at time t. Lemma For sello, $X_1^{(N)}(s+l_N) \gg X_N^{(N)}(s)$.

Warm up lemma Recall $l_N = \prod_{n=1}^{N} N^7$. Recall $X_{1}^{(N)}(t) \leq X_{2}^{(N)}(t) \leq ... \leq X_{N}^{(N)}(t)$ ordered particle positions at time t. Lemma For sello, $X_{1}^{(N)}(s+l_{N}) \gg X_{N}^{(N)}(s)$. Proof: Suppose (for a contradiction) that $X_{i}^{(N)}(t) < X_{N}^{(N)}(s)$ $\forall t \in [s, s + l_{N}] \cap \mathbb{N}_{0}$. Then since all jumps are non-negative, the rightmost particle at time s has 2^{l_N} descendants at time $s+l_N$. Since $2^{l_N} \ge N$, this implies $X_1^{(N)}(s+l_N) \ge X_N^{(N)}(s)$. So $\exists s^* \in [s, s+l_N] \cap \mathbb{N}_0 \ s.t. \ X_1^{(N)}(s^*) \geqslant X_N^{(N)}(s).$ All jumps are > 0, so $X_1^{(N)}(s+l_N) \ge X_1^{(N)}(s^*)$. \Box .

Construction of N-BRW from BRWs.

BRW: Initial particle at xER at time O.

At each time n E INo, each time-n particle has two offspring, each of which makes an independent jump from its parent's location with the same distribution as X. The time-(n+1) particles are these offspring particles.

Number of time-t particles is 2^t.

Construction of N-BRW from BRWs. BRW: Initial particle at xER at time O. At each time nENO, each time-n particle has two offspring, each of which makes an independent jump from its parent's location with the same distribution as X. The time-(n+1) particles are these offspring particles. Number of time-t particles is 2^t. Construction of N-BRW with initial particle locations $\infty_1, ..., \infty_N$: Take N independent BRWs with initial particles at $x_1, ..., x_N$. Colour BRW particles blue or red. All time-O particles are blue. For nello, the N rightmost offspring particles of time-n blue particles are coloured blue. All other time-(n+1) particles are coloured red. Blue particles form an N-BRW. Path of jumps from ancestor to descendant in N-BRW = path in one of the BRWs. Random walk with heavy tailed jump distribution $P(X > \infty) = \frac{1}{h(\infty)}$, h regularly varying with index $\alpha > 0$. $X_{1}, X_{2}, X_{3}, \dots$ i.i.d. with $X_{1} \stackrel{d}{=} X$.

Fix c>O small. For ∞ v. large, unlikely that $\sum_{k=1}^{n} X_k \ge \infty$ and $X_k \le c\infty$ $\forall k \le n$.

Random walk with heavy tailed jump distribution $\mathbb{P}(X > \infty) = \frac{1}{h(\infty)}$, h regularly varying with index $\infty > 0$. X_1, X_2, X_3, \dots i.i.d. with $X_1 \stackrel{d}{=} X_1$ Fix c>O small. For ∞ v. large, unlikely that $\sum_{k=1}^{\infty} X_k \ge \infty$ and $X_{k} \leq c\infty \quad \forall k \leq n.$ Lemma (Durrett '83, Gantert '00) For mEN, q>0, $\lambda>0$ and $0 < r < 1 \land \frac{\lambda(1 \land \alpha)}{8q}$, for N sufficiently large, if $\mathbb{P}\left(\sum_{i=1}^{m\ell_{N}} X_{j} \mathbf{1}_{X_{j} \leq r \times_{N}} \ge \infty_{N}\right) \le N^{-9}.$ $x_N > N^{\lambda}$ then

Random walk with heavy tailed jump distribution Lemma (Durrett '83, Gantert '00) For mEN, q>0, λ >0 and $0 < r < 1 \land \frac{\lambda(1 \land \alpha)}{8q}$, for N sufficiently large, if $x_N > N^{\lambda}$ then $P(\sum_{j=1}^{mN} X_j 1 = x_N \ge x_N) \le N^{-q}$.

Random walk with heavy tailed jump distribution Lemma (Durrett '83, Gantert '00) For mEN, q>0, λ >0 and $0 < r < 1 \land \frac{\lambda(1 \land \alpha)}{8q}$, for N sufficiently large, if $\mathbb{P}\left(\sum_{i=1}^{m\ell_{N}} X_{j} \mathbb{1}_{X_{j}} \leq r x_{N} \geq x_{N}\right) \leq N^{-9}.$ $x_N > N^{\lambda}$ then Proof: Take $2q \log 2 < c < \frac{\lambda(1 \wedge \alpha) \log 2}{-}$ N^{c/log2} By Markov's inequality, $\mathbb{P}\left(\sum_{j=1}^{m\ell_{N}} X_{j} \mathcal{I}_{X_{j} \leq r \alpha_{N}} \geqslant \alpha_{N}\right) = \mathbb{P}\left(e^{c\ell_{N} \alpha_{N}^{-1}} \sum_{j=1}^{m\ell_{N}} X_{j} \mathcal{I}_{X_{j} \leq r \alpha_{N}} \geqslant e^{c\ell_{N}}\right)$ $\leq N^{-\frac{c}{\log 2}} \mathbb{E}\left[e^{c\ell_{N} \alpha_{N}^{-1}} X \mathcal{I}_{X \leq r \alpha_{N}}\right]^{m\ell_{N}}.$ Use identity $\mathbb{E}\left[e^{v \mathcal{Y} \mathbf{1}}_{\mathcal{Y} \neq K_{2}} \mathbf{1}_{\mathcal{Y} \neq K_{1}}\right] = \int_{V}^{K_{2}} v e^{v u} \mathbb{P}(\mathcal{Y} \neq u) du + e^{v K_{1}} \mathbb{P}(\mathcal{Y} \neq K_{1}) - (e^{v K_{2}} - 1) \mathbb{P}(\mathcal{Y} \neq K_{2})$ to show $\mathbb{E}\left[e^{c\ell_N x_N^{-1} \times 1 \times (x - rx_N)}\right] = 1 + \mathcal{O}(N^{-\epsilon})$ for some $\epsilon > 0$.

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Random walk with heavy tailed jump distribution Lemma (Durrett '83, Gantert '00) For mEN, q>0, $\lambda>0$ and $0 < r < 1 \land \frac{\lambda(1 \land \alpha)}{8q}$, for N sufficiently large, if $x_N > N^{\lambda}$ then $\mathbb{P}\left(\sum_{j=1}^{N} X_{j} \mathbb{1}_{X_{j} \leq r \times_{N}} \ge x_{N}\right) \le N^{-9}.$ Use with $x_N = const \cdot a_N$ Recall $a_N = h^{-1}(2N\ell_N)$, so $a_N'' \sim (2N\ell_N)'^{\prime \alpha}$. Fix $e \in (0, 1)$ small. A jump larger than ea_N is a "big jump". For $c \gg e$, it is very unlikely that there is a time-t particle > ca_N away from its time - $(t - O(\ell_N))$ ancestor unless an ancestor made a big jump. A jump is big $\omega.p.$ $h(ea_N)^{-1} \sim e^{-\alpha} (2N\ell_N)^{-1}$, So we see big jumps at rate $\oplus(l_N^{-1})$.





A: A particle makes a big jump at time T and takes the lead (by $\Theta(a_N)$). Its descendants stay in the lead until time t_1 (other particles can't take the lead with a big jump, and can't move far without a big jump).



B: There are O(1) big jumps in time interval $[t_1, t]$, each with O(N) descendants at time t.



C: The tribe descended from the time-T leader doubles in size at each timestep until almost time $T + l_N$.



D: On the time interval $[T+l_N,t]$, the time-T leader's tribe has size N-O(N).



 $X_{i}^{(N)}(s+l_N) \ge X_{N}^{(N)}(s)$ and so $\frac{1}{a_N}(X_{N}^{(N)}(s+l_N) - X_{i}^{(N)}(s+l_N))$ is small $\omega.h.p.$





Star-shaped genealogy No time- $(T + \varepsilon_N l_N)$ particles have $\Theta(N)$ time-t descendants. None of the particles in the time-T leader's tribe have moved far by time $T + \varepsilon_N l_N$, so each has $\Theta(N 2^{-\varepsilon_N l_N}) = o(N)$ descendants at time t.